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# Pulsed Lidar for the ASCENDS Mission: Space Instrument Studies

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# ASCENDS Mission



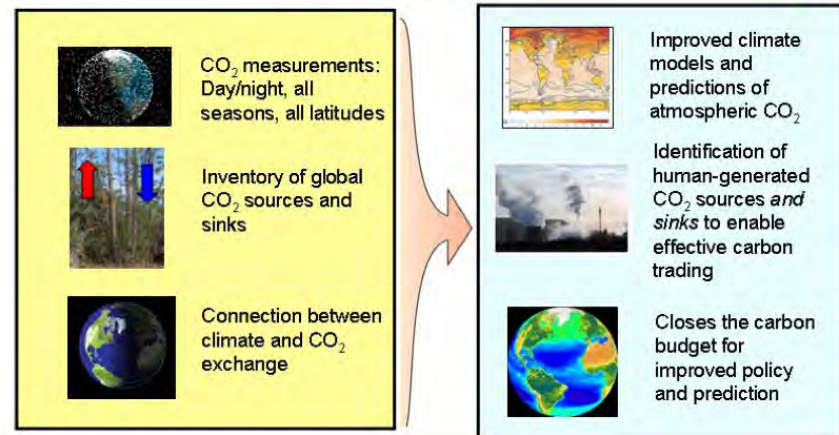
## Active Sensing of CO<sub>2</sub> Emissions over Nights, Days, and Seasons (ASCENDS)

### Why lasers ?

- Measures at night & all times of day
- Constant nadir/zenith path
  - Illumination = observation path
  - Continuous "glint" measurements over oceans
- Measurements at high latitudes
- Small measurement footprint
- Measure through broken clouds
- Measure to cloud tops
- Very high spectral resolution and accuracy

Our approach →

Active Sensing of CO<sub>2</sub> Emissions over Nights, Days, and Seasons (ASCENDS)  
 Launch: 2013-2016  
 Mission Size: Medium



### Lidar approaches for CO<sub>2</sub> column:

- Broadband laser - 1570 nm band -  $\lambda$  tuned receiver
- 1 line - 2 um band - pulsed - direct detection
- 1 line - 2 um band - CW heterodyne detection
- 1 line - 1570 nm band - synchronous direct detection
- 1 line - 1570 nm band - pulsed direct detection



# Laser Sounder Approach for ASCENDS Mission



## 3 simultaneous laser measurements

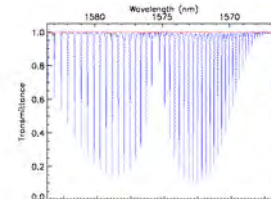
1. CO2 lower tropospheric column  
One line near 1572 nm
2. O2 total column (pressure)  
Measured between 2 lines near 765 nm
3. Altimetry & atmospheric backscatter profile from CO2 signal  
Surface height and atmospheric scattering profile at 1572 nm

### Measurements use:

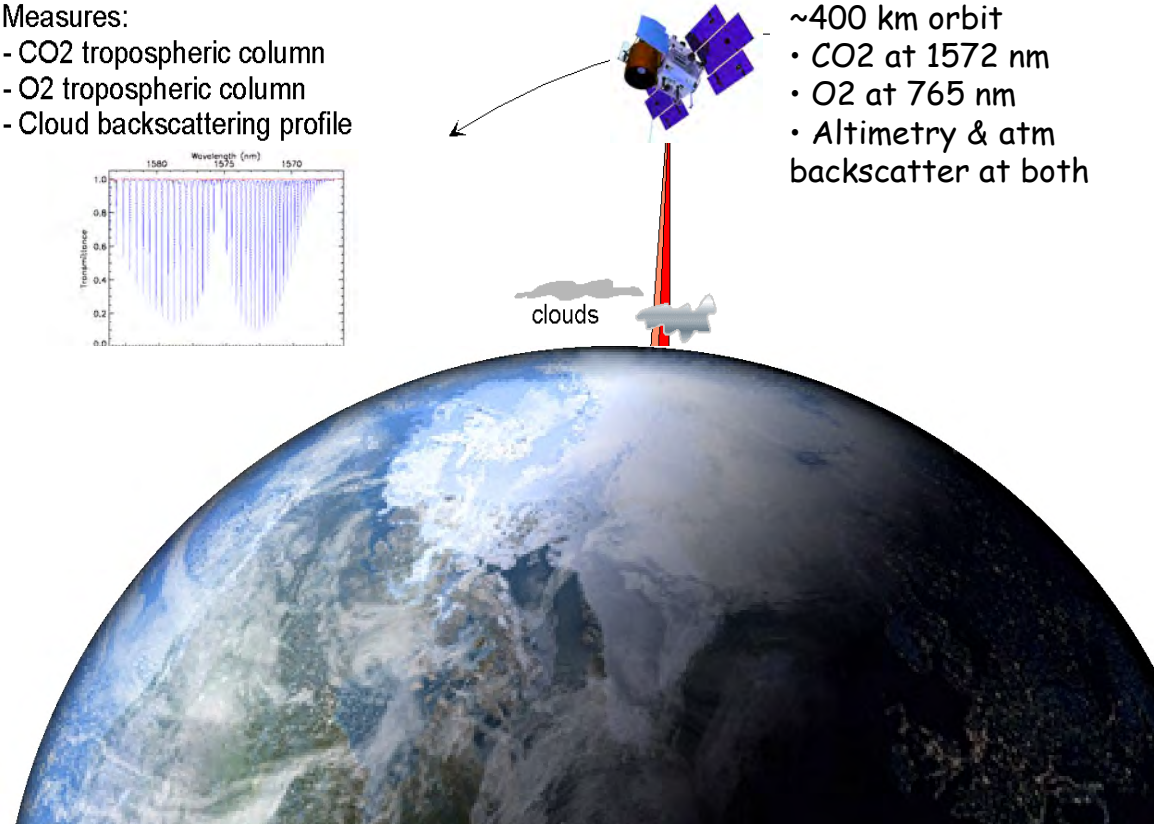
- Pulsed EDFA lasers
- 8 KHZ pulse rates
- ~ 8 laser wavelengths/ gas line
- Time gated Photon counting receiver

### Measures:

- CO2 tropospheric column
- O2 tropospheric column
- Cloud backscattering profile



- ~400 km orbit
- CO2 at 1572 nm
- O2 at 765 nm
- Altimetry & atm backscatter at both



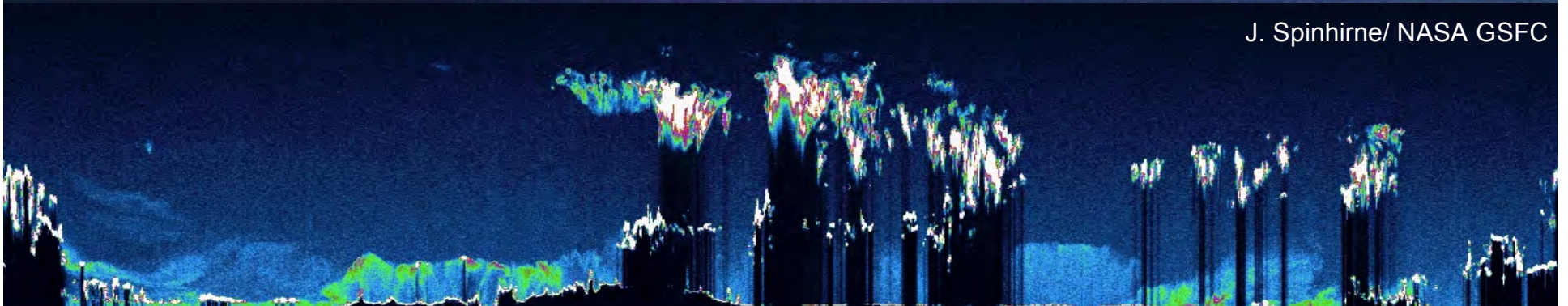
### CO2 & O2 column measurements:

- Pulsed (time gated) signals :
  - Isolates full column signal from surface
  - Reduces noise from detector & solar background
- Target: ~ 1ppmV in ~100 km along track sample

**Atmospheric Scattering is widespread  
& is quite complex**  
**Pulsed Lidar can measure to the surface through  
scattering**

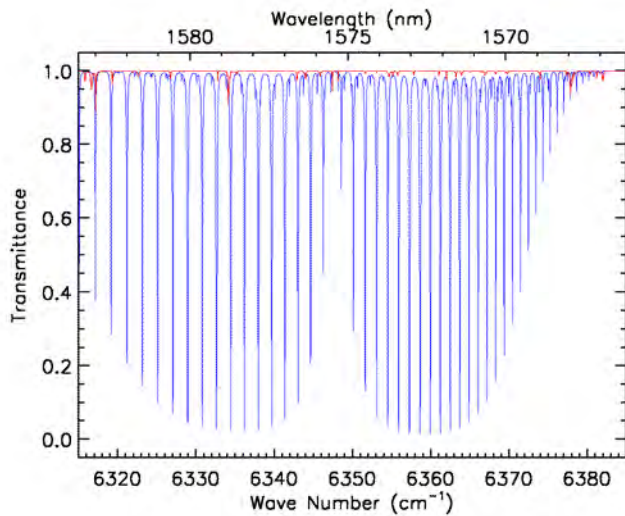
The vertical dimension:  
The Geoscience Laser Altimeter System – 2003  
Aerosol and Cloud Laser Measurements from Space

J. Spinhire/ NASA GSFC

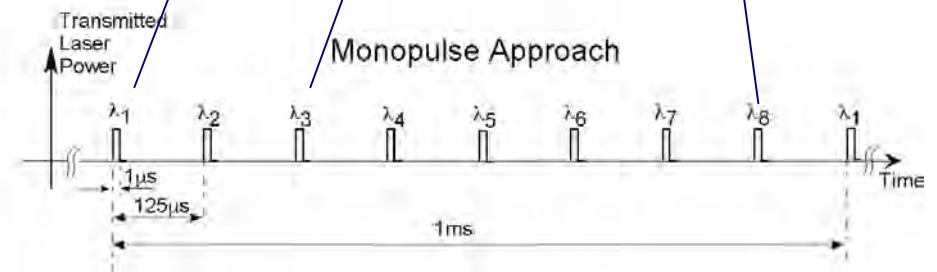
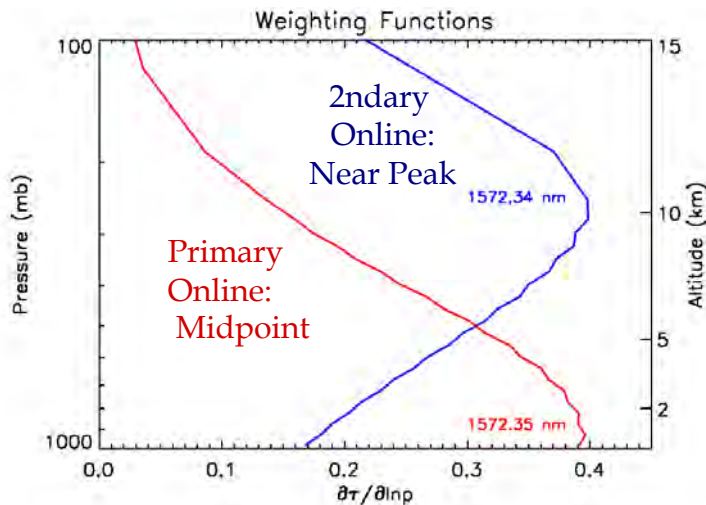
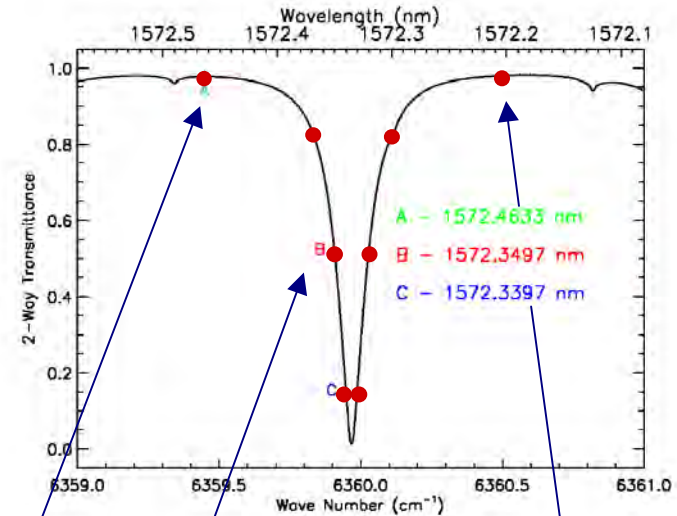




# Candidate CO2 Line, Sampling & Vertical Weighting Functions



Space lidar plan is ~8 wavelength samples across line

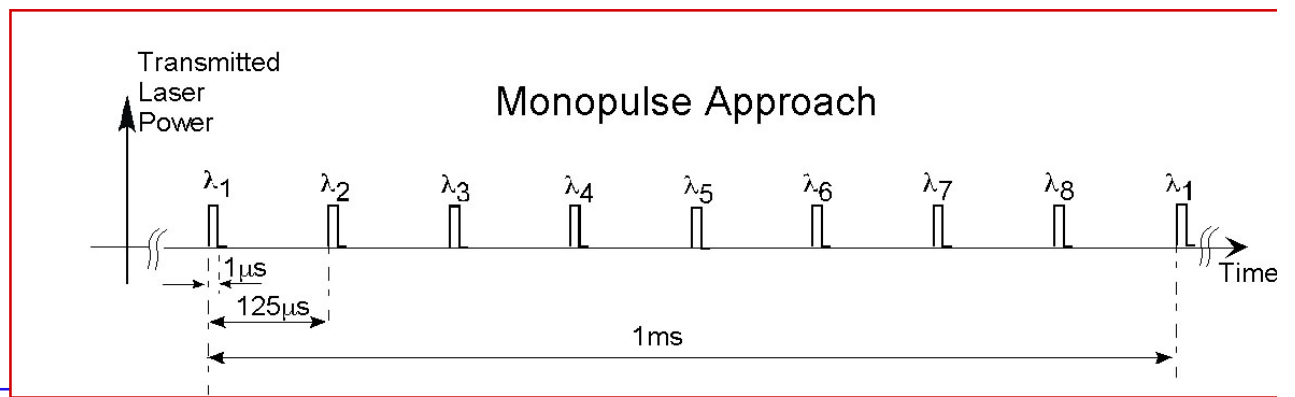
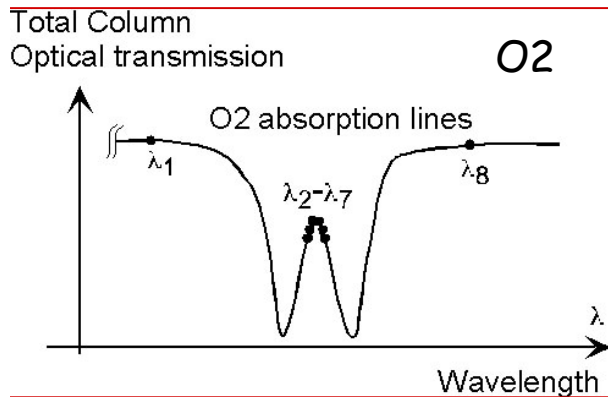
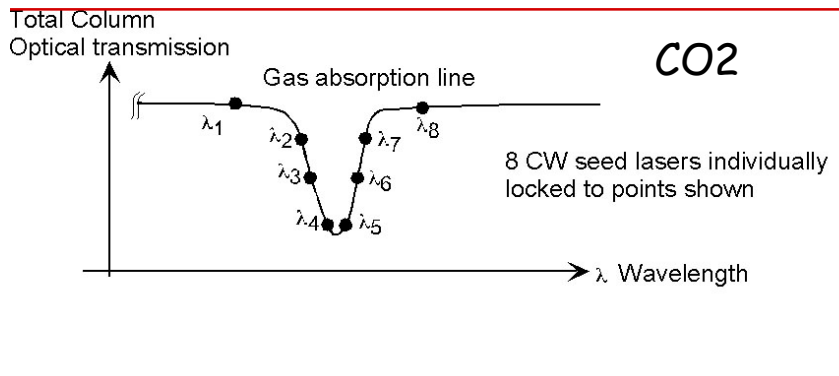
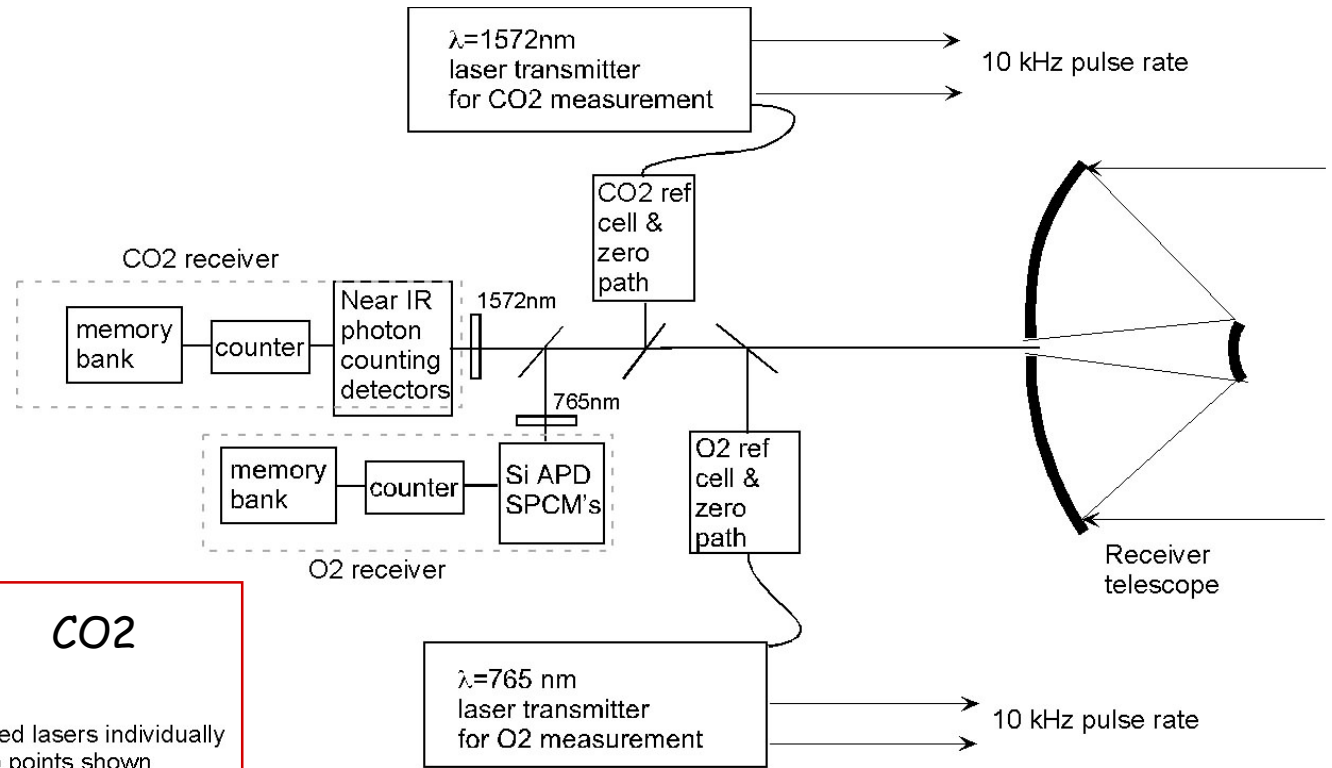


Multi-wavelength Line Sampling allows:

- Detection & correction of Doppler & λ errors
- Modeling -> reducing errors from varying λ response
- CO2 retrievals for: Lower troposphere
- Total column; Line shape information

Total Column from line area

# Measurement approach for Space





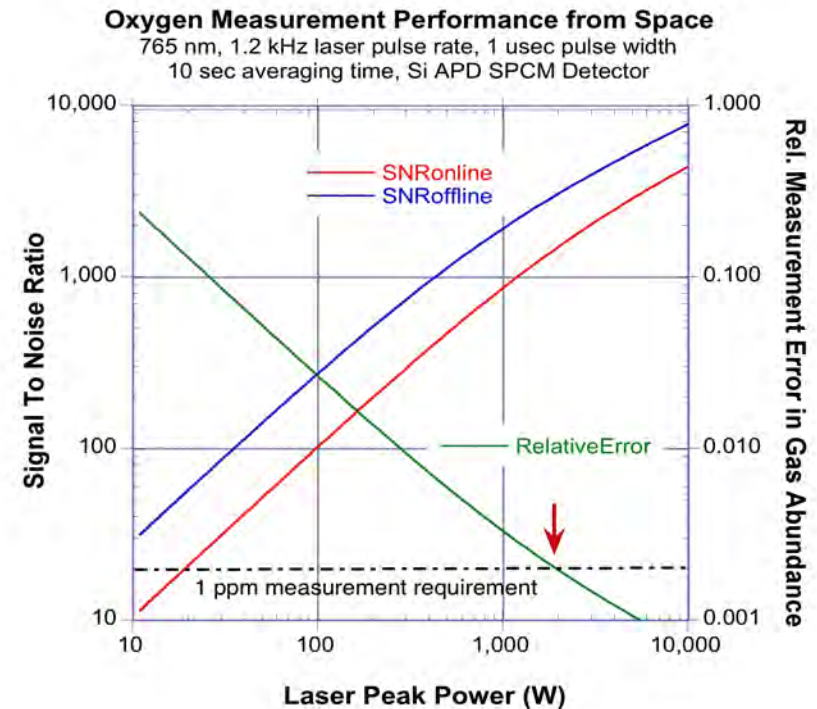
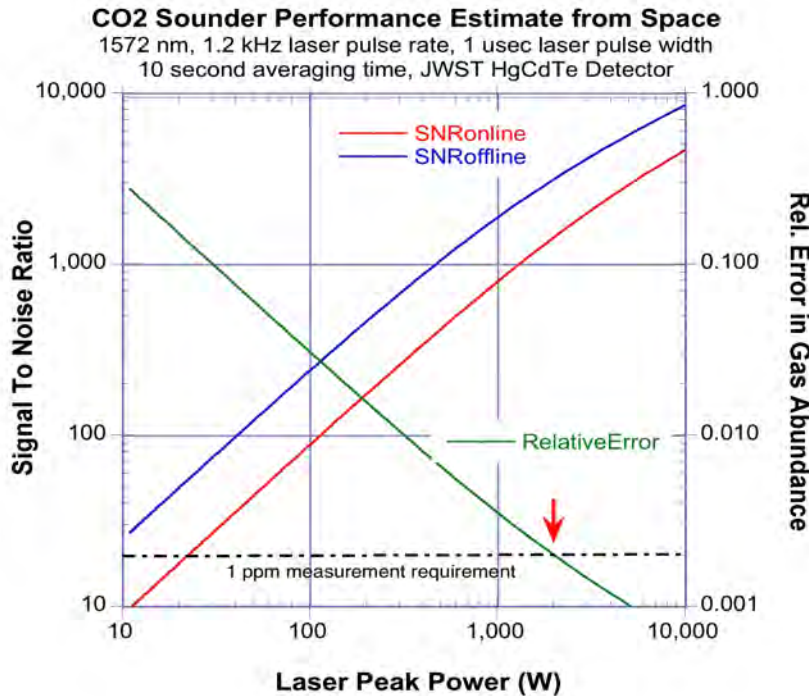
# Space: SNR & Relative Measurement Errors

(10 seconds observing time, 500\* km orbit, 1.5m telescope)



## CO2 column measurement

## O2 column measurement



~ 3 mJ/pulse energy (PMT detector)

~ 3 mJ/pulse energy (PMT detector)

Ave optical power ~25-30W

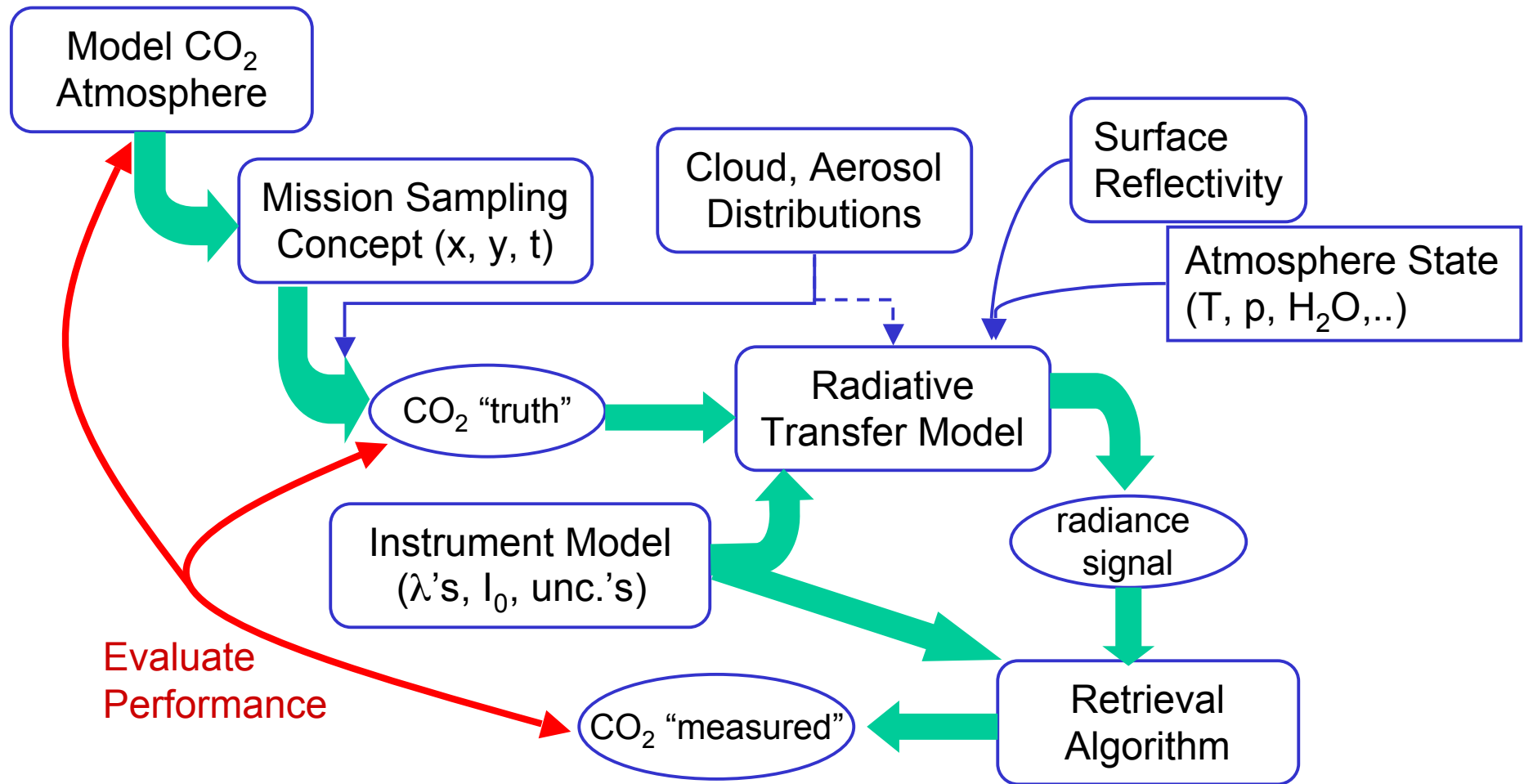
6 mJ energy from 1530 nm amp, 50% doubling

Ave optical power ~25-30 W

\* - Same performance at 3 mJ/pulse with PMTs at 400 km orbit



# Measurement Model used in Mission Performance Simulation (Randy Kawa)

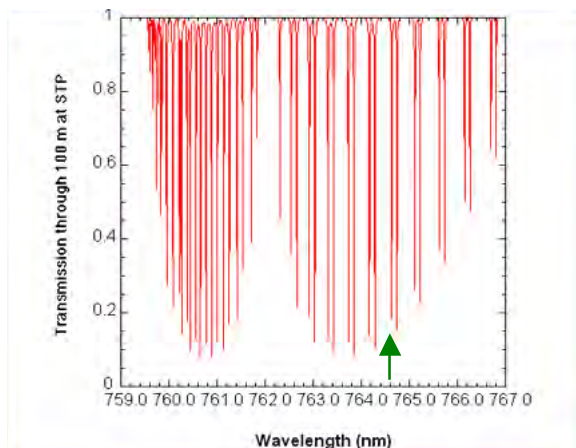


→ Test sensitivity of inferred CO<sub>2</sub> distributions to varying mission and instrument design parameters.



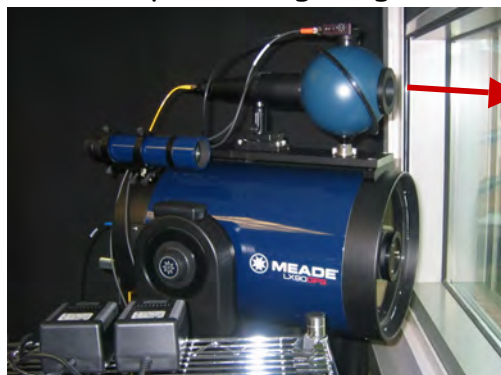


# Oxygen - Open path measurement of absorption lines near 765 nm (M. Stephen)

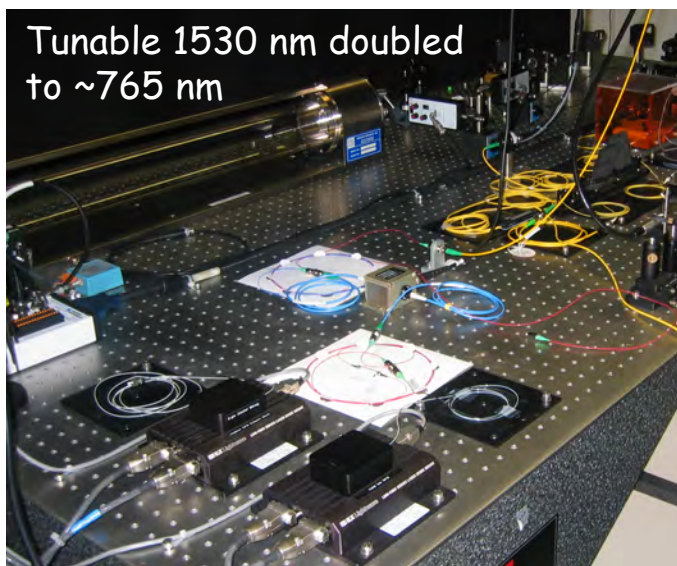
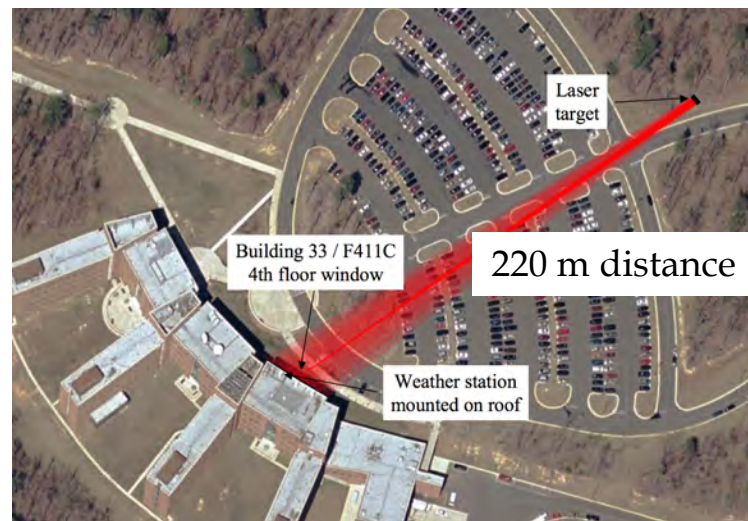


Oxygen A band: Calculated atmospheric transmission for 100 m path at STP

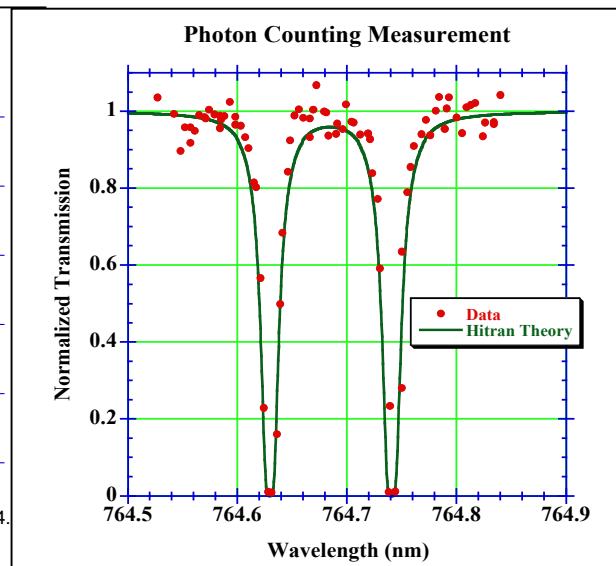
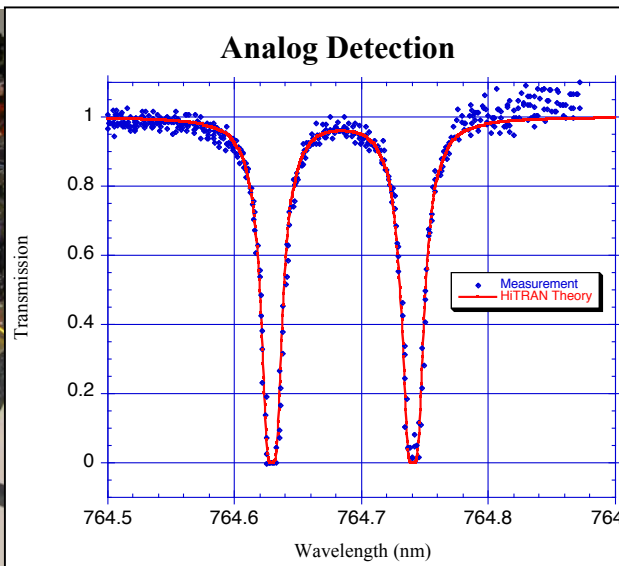
Telescope viewing target



Peak optical power ~ 50 mW  
Attenuation for round trip was  $\sim 10^6$



Tunable 1530 nm doubled to  $\sim 765$  nm

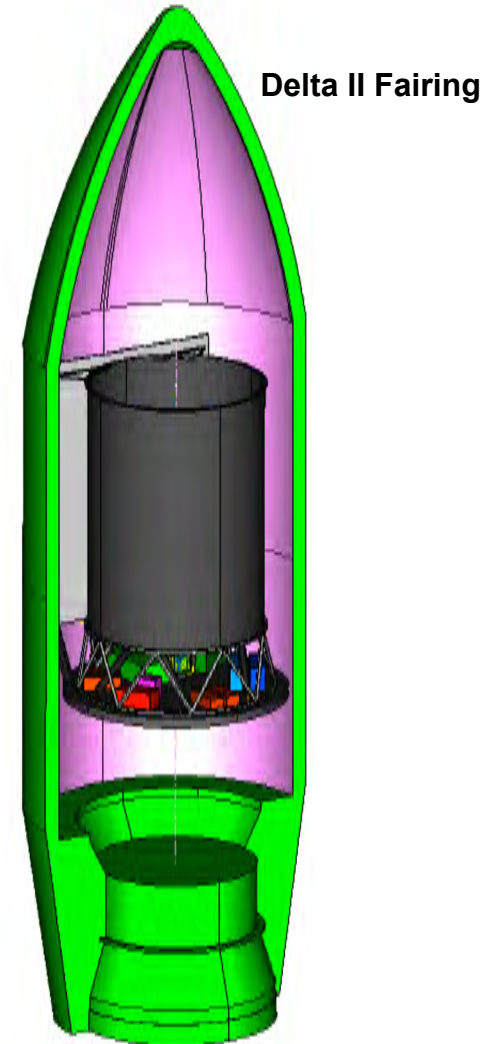




# Initial Space Instrument Study

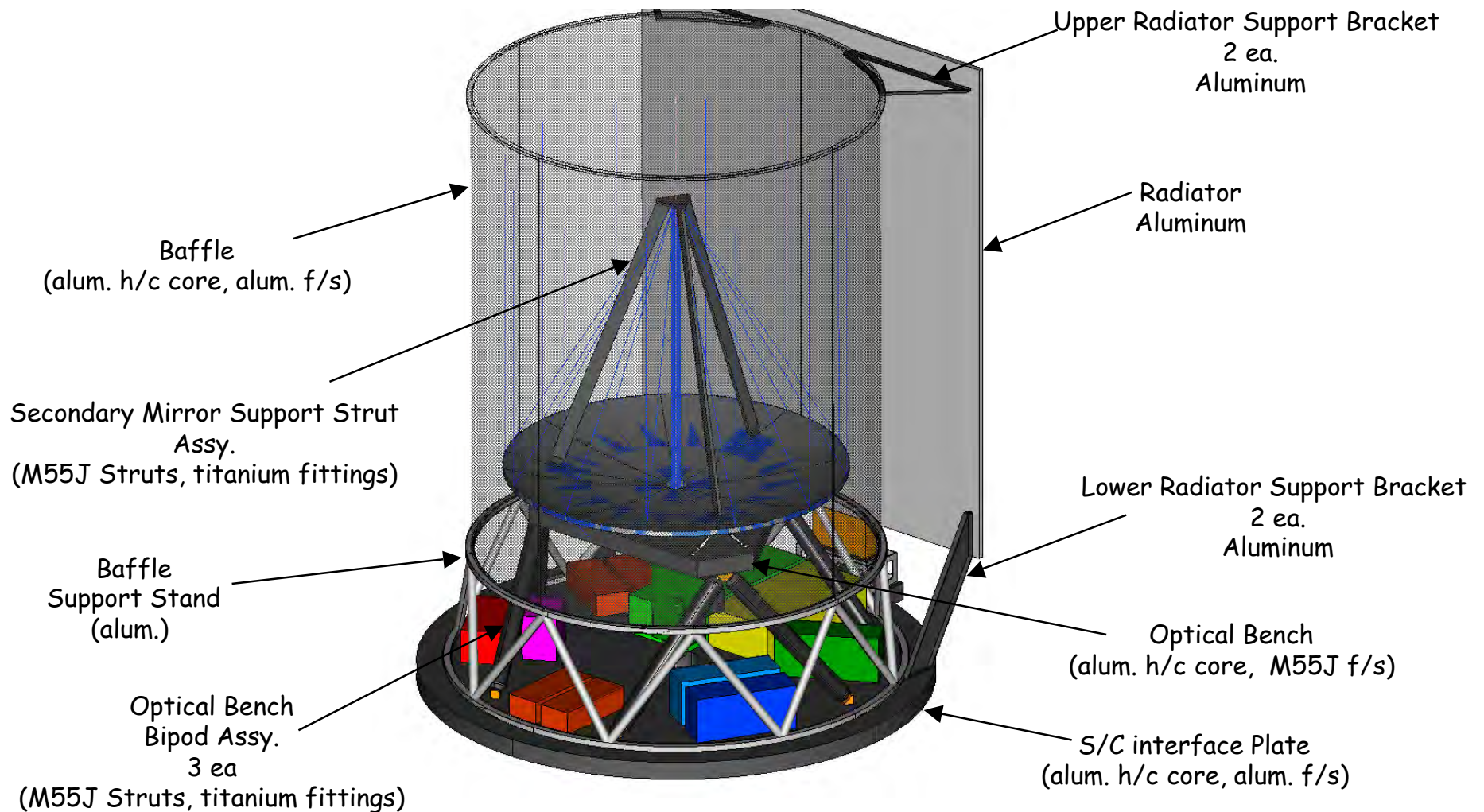


- Sun synchronous orbit
  - Altitude 500 km
  - Sun-sync inclination, 1:30 pm crossing time
- Mission Risk Class B
  - 5 year mission life
  - 85% mission reliability
  - Mitigate single point failures with redundancy or high reliability parts
- Traditional S/C bus orbit and attitude knowledge sufficient for mission requirements (i.e., no on-instrument attitude processing required)





# Initial Concept Layout for Lidar (telescope diameter = 1.5 m)

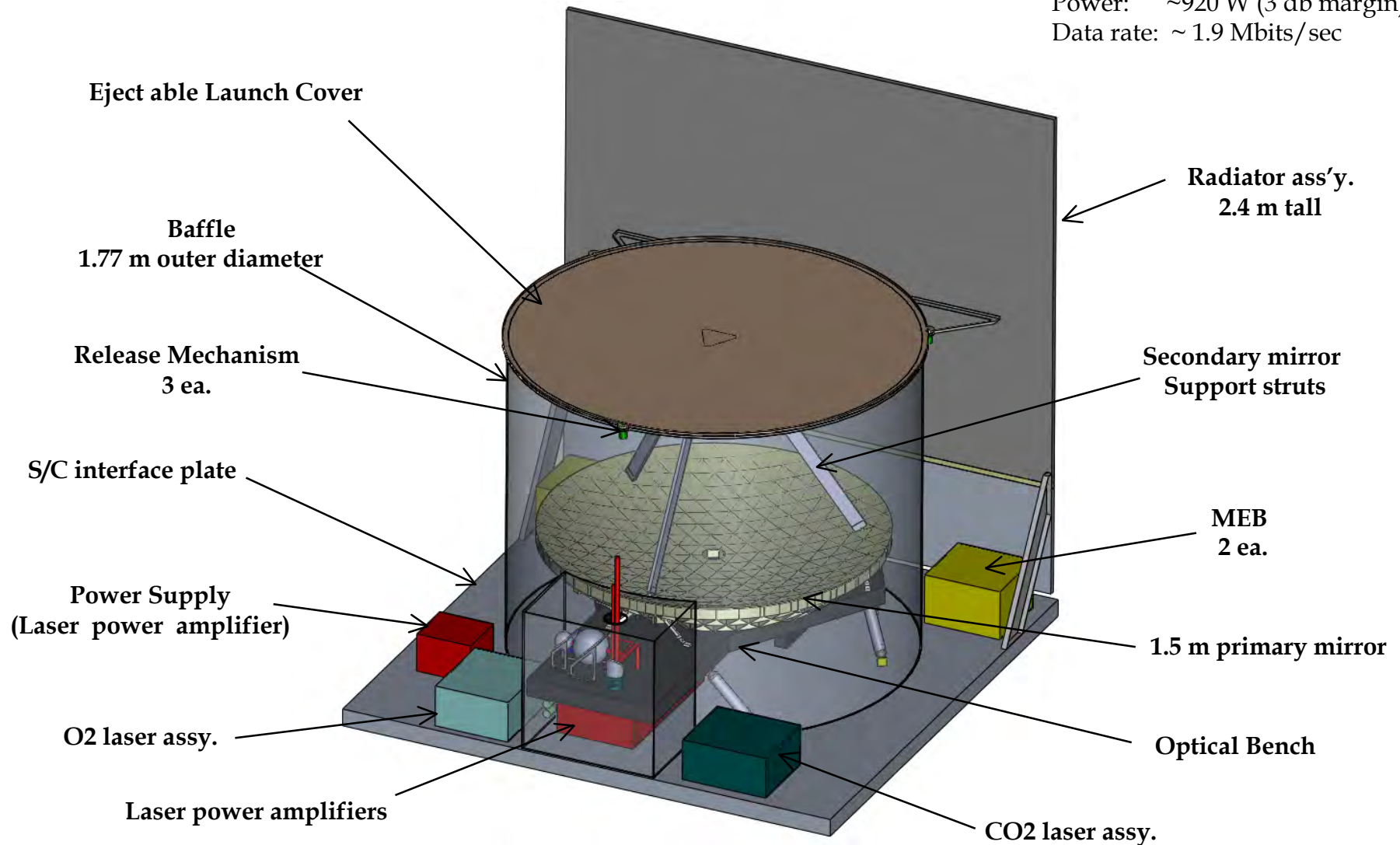




# Second Study: ASCENDS Lidar Configuration (layout not optimized)

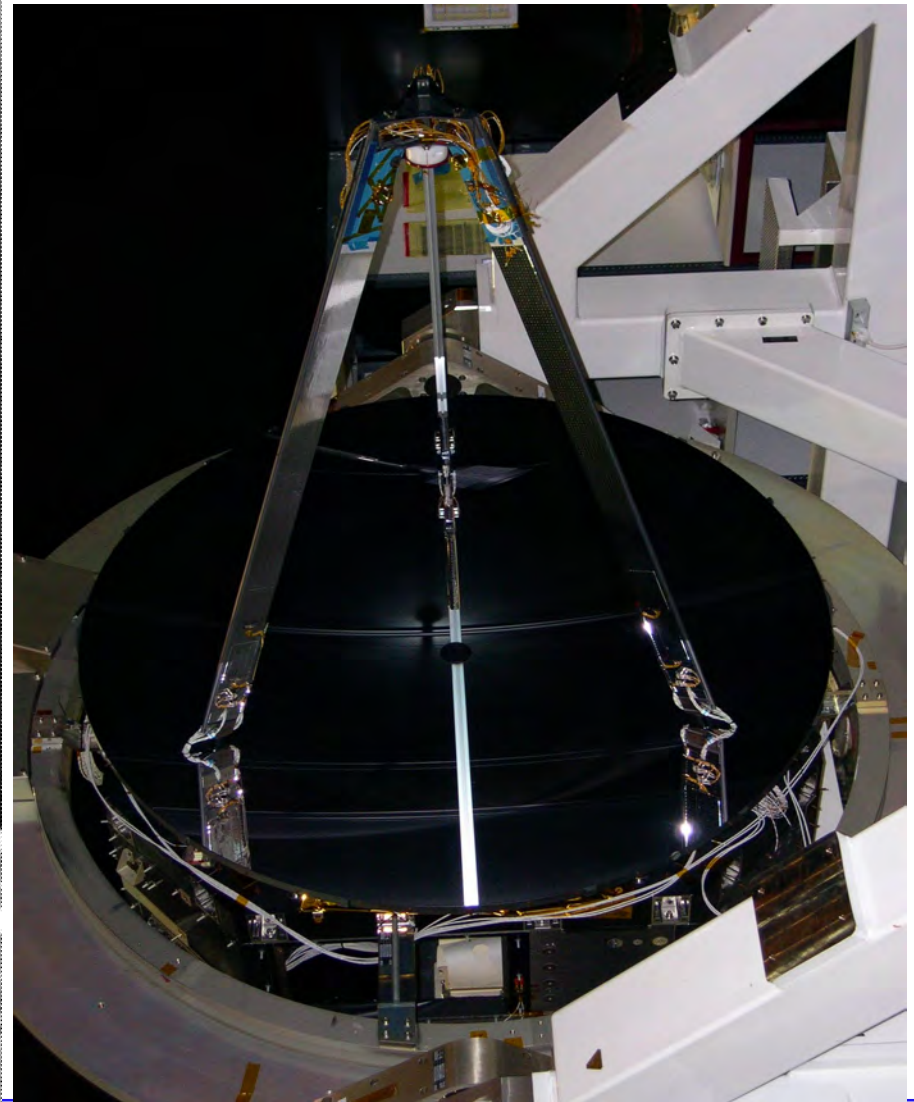
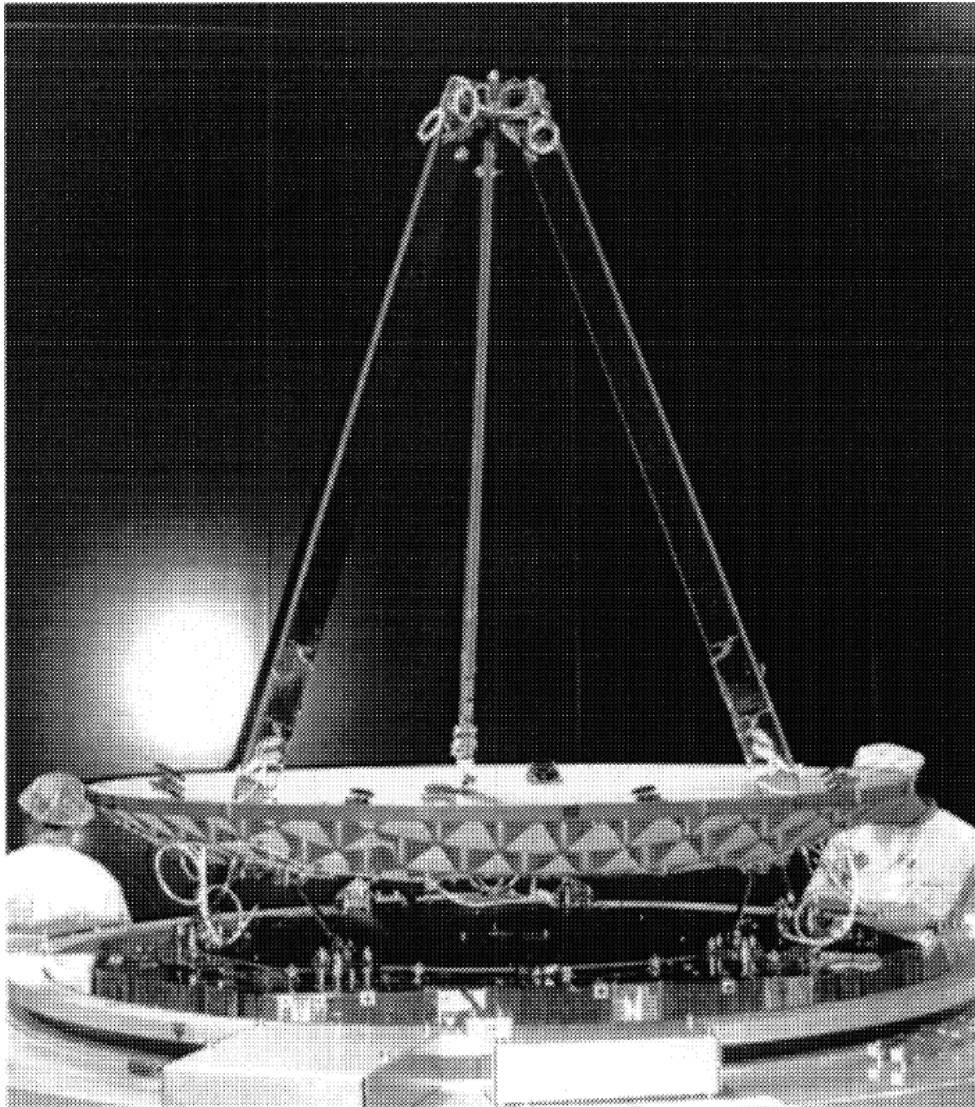


Mass: ~420 kg  
Power: ~920 W (3 db margin)  
Data rate: ~1.9 Mbits/sec





# ESA ALADIN 1.5 m Receiver Telescope



3-15-10

NASA TE Meeting: Pulsed Lidar for ASCENDS Mission: Space Instrument Studies

Abshire - 13



# CO<sub>2</sub> Receiver - Photon Counting Detectors



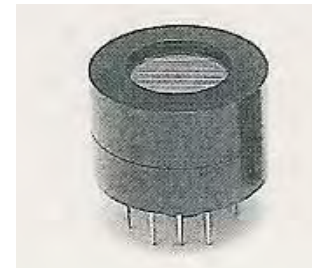
HV supply and PMT housing

*Now:*

## Hamamatsu H9170-75 PMT

-Detector used in airborne receiver

- Turn-key operation
- QE = 2-16% at 1570nm
- InP/InGaAs photocathode
- Photocathode: ~5 mm diameter
- Dark count rate ~200 KHz at -80 C (TEC cooled)
- PMT power consumption ~150mW



TO-8 PMT package with transmissive photocathode

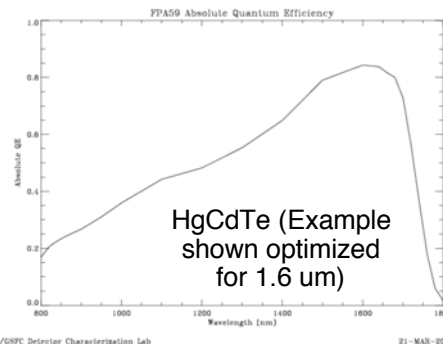


## The Infrared Detectors for the Wide Field Camera 3 on HST

### ABSTRACT

We present the performance of the IR detectors developed for the WFC3 project. These are HgCdTe 1Kx1K devices with cutoff wavelength at 1.7 $\mu$ m and 150K operating temperature. The two selected flight parts, FPA#64 (prime) and FPA#59 (spare) show quantum efficiency higher than 80% at  $\lambda=1.6\mu$ m and greater than 40% at  $\lambda>1.1\mu$ m, readout noise of ~25 e<sup>-</sup> rms with double correlated sampling, and mean dark current of ~0.04 e<sup>-</sup>/s/pix at 150K. We also report the results obtained at NASA GSFC/DCL on these and other similar devices in what concerns the QE long-term stability, intra-pixel response, and dark current variation following illumination or reset.

## Under Development: Adapting HgCdTe Photon Counting Detectors



80% photon detection/accumulation efficiency at 1550 nm with ~0.04 dark counts/sec at 150K

Present read-out configured for photon accumulation



## O<sub>2</sub> Channel Detectors: Ones used for 532 nm channel of ICESat/GLAS Lidar



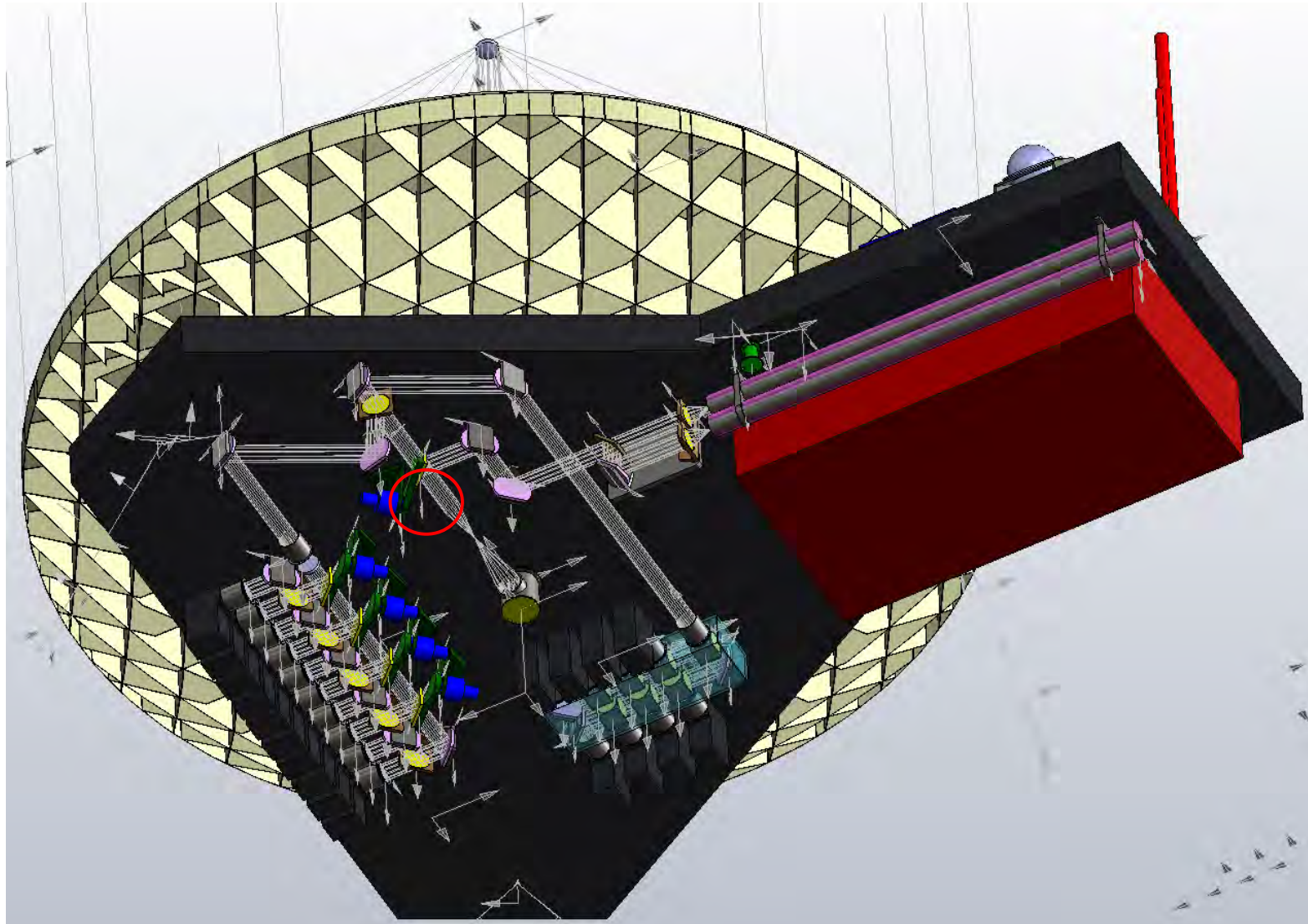
### O<sub>2</sub> Channel Detectors

- SCPM Detectors (8 each) From Perkin Elmer
- ~70% QE (765 nm) Now TRL=9
- Successfully improved & up-screened commercial units for GLAS/ICESat
- Advantages: High Sensitivity High TRL





## 2nd study - Aft view of optical bench







# 2nd ASCENDS Lidar Study: Two candidate rockets (both have considerable excess lift capacity)



## Space-X Falcon 9

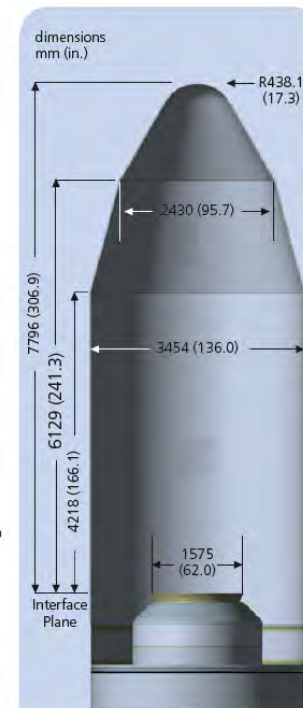


Quarter section of the 5.2 m Falcon 9 fairing at SpaceX's Hawthorne, CA headquarters.

**Ascends Observatory (lidar + spacecraft + fuel) needs only ~1/4 lift capacity**



Falcon 9 Engines Close Up

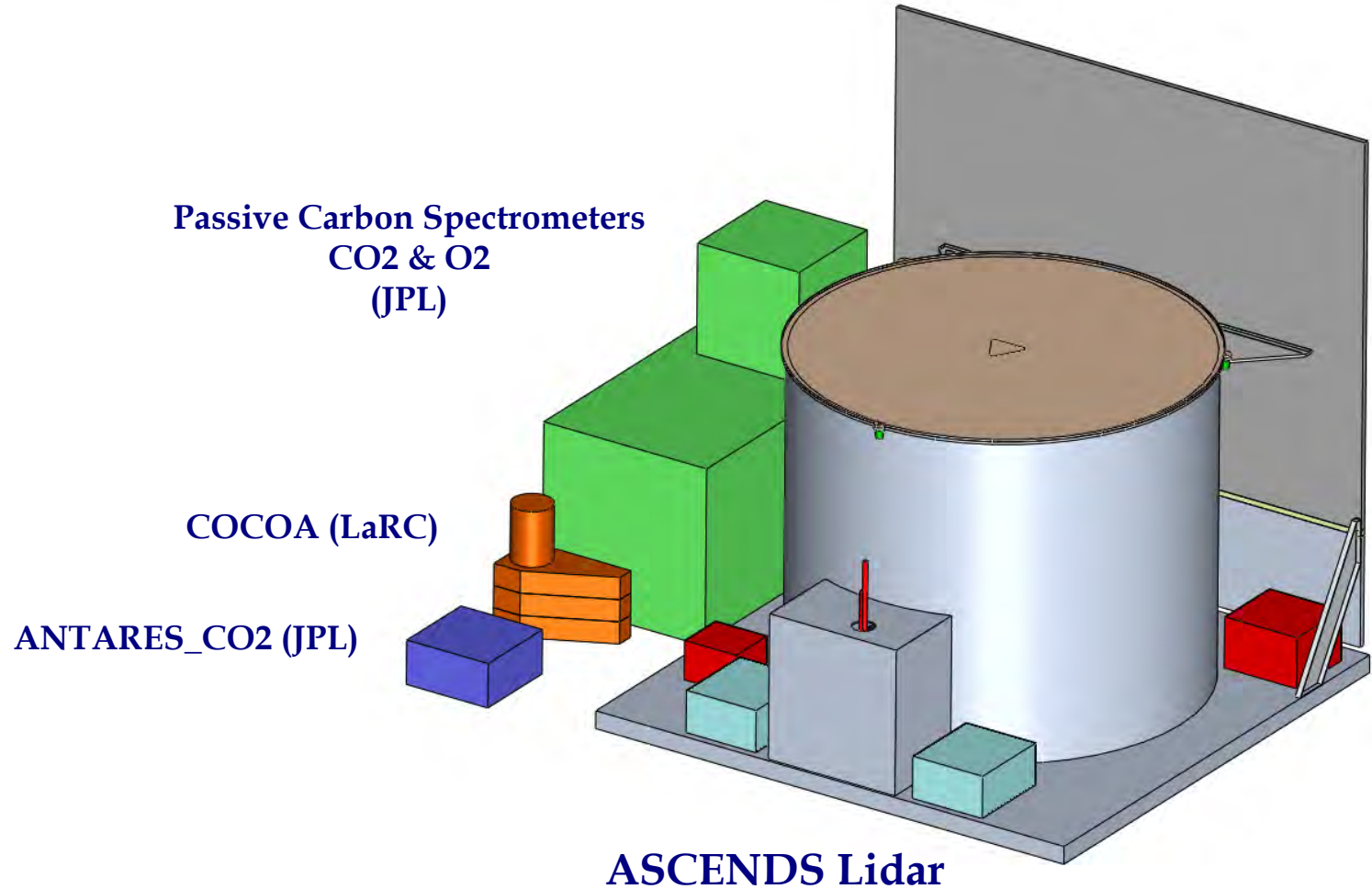


**Taurus II** Medium-Class Launch Vehicle  
INNOVATION YOU CAN COUNT ON

**Ascends Observatory needs ~1/2 capacity**

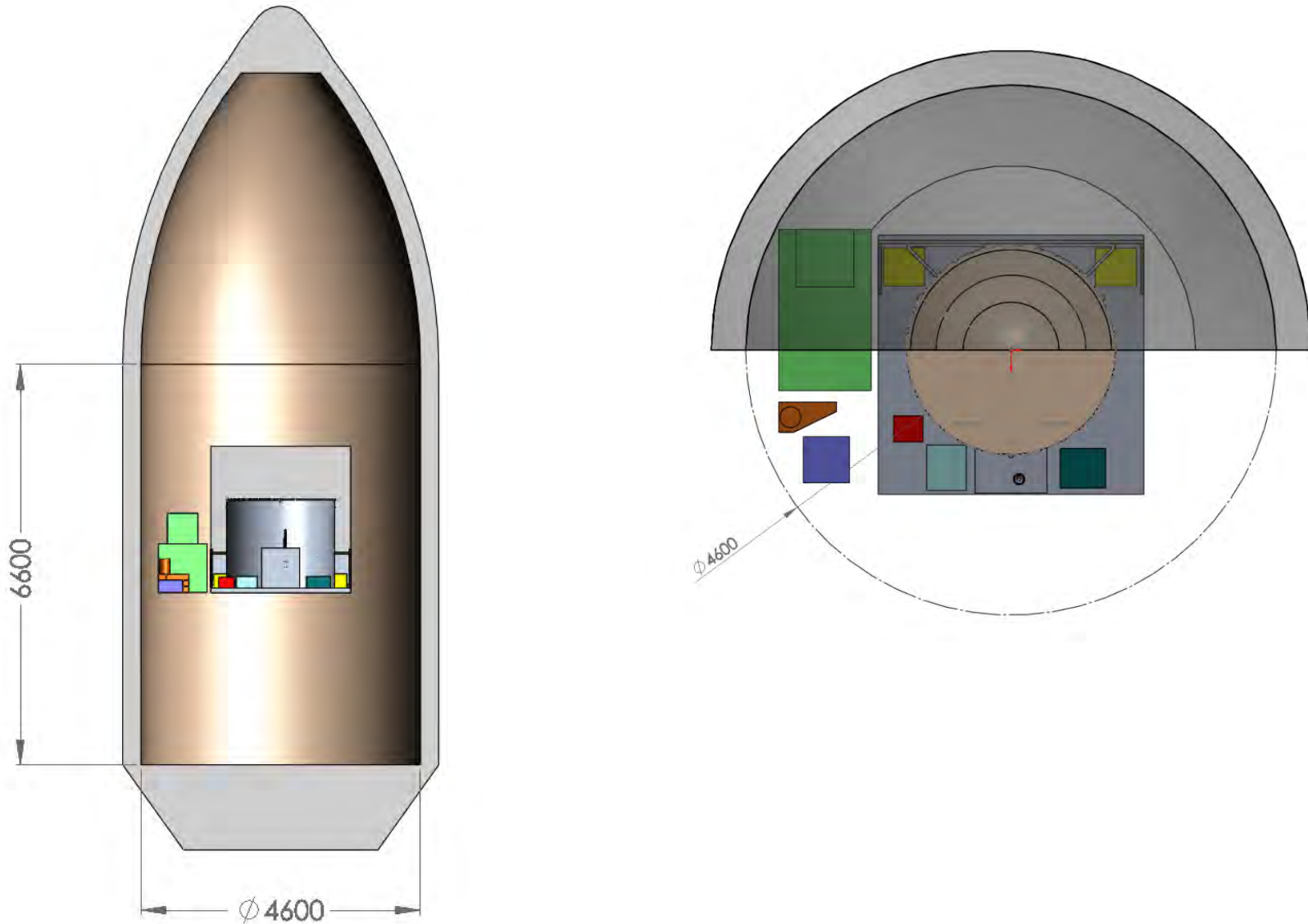


# 2nd Study: ASCENDS lidar and some candidate passive spectrometers





# ASCENDS Lidar & 3 Spectrometers in a Falcon 9 Faring





# ASCENDS Space Lidar Studies - Summary



- Conducted two studies of this approach for ASCENDS space lidar in 4/08 and 9/09
- Measures CO<sub>2</sub> & O<sub>2</sub> via pulsed direct detection from ~ 400 km, sun-sync 1:30 pm orbit
- Straightforward space lidar design:
  - Mass: ~ 420 Kg (can be reduced via more efficient layout)
  - Power: ~ 800 - 920W (3dB margin); driven by SNR needs, detector choice
  - Data rate: ~1.9 Mbit/sec - Ok for use of high latitude comm. ground site
  - Size - driven by receiver telescope diameter and radiators needed for lasers
- Low risk: use of space qualified telescope, O<sub>2</sub> detectors, thermal system
- Primary power draw & thermal source: laser subassemblies
- Laser architecture -> high efficiency & reliability, switchable cold spares
- Detectors -> reliability via use of multiple detectors & switchable spares.
- Several aspects can be optimized in follow on studies